

Constrained SUSY after the Higgs discovery

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The status of the CMSSM



healthy?



pretty dull?



almost dead?



buried?

Fittino

C++ program for model testing and parameter analysis using

- ▶ all kind of available experimental measurements
- ▶ public theory codes to calculate predictions
- ▶ a χ^2 function to compare measurements and predictions
- ▶ an auto-adaptive Markov Chain to sample the parameter space
- ▶ frequentist interpretation

Publications:

[arXiv:0412012 \[hep-ph\]](#), [arXiv:0511006 \[hep-ph\]](#), [arXiv:0907.2589 \[hep-ph\]](#),
[arXiv:0909.1820 \[hep-ph\]](#), [arXiv:1102.4693 \[hep-ph\]](#),
[arXiv:1105.5398 \[hep-ph\]](#), [1204.4199 \[hep-ph\]](#), [arXiv:1310.3045 \[hep-ph\]](#)

The CMSSM

Constrained MSSM reduces 124 parameters of MSSM to 4 and a sign:

M_0 common scalar mass parameter

$M_{1/2}$ common gaugino mass parameter

A_0 common trilinear coupling

$\tan \beta$ ratio of Higgs VEVs

$\text{sgn } \mu$ sign of Higgsino mass parameter

We fix $\text{sgn } \mu = +1$

We use m_t as additional free fit parameter.

Fitting the CMSSM

CMSSM is experimentally constrained by

- ▶ indirect constraints from low energy precision measurements
- ▶ SM-like Higgs boson
- ▶ direct searches for sparticles and Higgs bosons
- ▶ astrophysical observations

To evaluate the corresponding model predictions we use:

- ▶ **SPheno** for spectrum calculation
- ▶ **FeynHiggs** for Higgs properties, $(g - 2)_\mu$ & Δm_s
- ▶ **SuperIso** for other B-Physics observables
- ▶ **Prospino**, **Herwig++**, **Delphes** for direct sparticle searches
- ▶ **MicrOMEGAs** for dark matter relic density
- ▶ **DarkSUSY** via **Astrofit** for direct detection cross section

Measurements

Low energy observables

$\mathcal{B}(B_s \rightarrow \mu^+ \mu^-)$	$(3.20 \pm 1.50 \pm 0.76) \times 10^{-9}$
$\mathcal{B}(B^\pm \rightarrow \tau^\pm \nu)$	$(0.72 \pm 0.27 \pm 0.11 \pm 0.07) \times 10^{-4}$
$\mathcal{B}(b \rightarrow s \gamma)$	$(3.43 \pm 0.21 \pm 0.07 \pm 0.23) \times 10^{-4}$
Δm_s	$(17.719 \pm 0.043 \pm 4.200) \text{ ps}^{-1}$
$a_\mu - a_\mu^{\text{SM}}$	$(28.7 \pm 8.0 \pm 2.0) \times 10^{-10}$
m_W	$80.385 \pm 0.015 \pm 0.010$
m_t	$(173.18 \pm 0.94) \text{ GeV}$
$\sin^2 \theta_{\text{eff}}$	0.23113 ± 0.00021

SM-like Higgs boson

- ▶ Higgs signals via **HiggsSignals**

Direct searches for sparticles and Higgs Bosons

- ▶ Higgs limits via **HiggsBounds**
- ▶ LEP chargino mass limit
- ▶ ATLAS MET + jets + 0 lepton search (20fb^{-1})

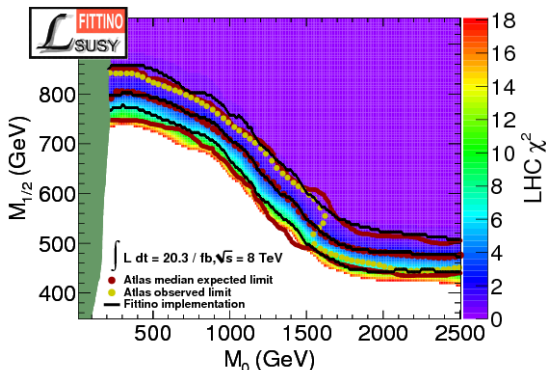
Astrophysical observables

- ▶ We require χ_1^0 to be the LSP
- ▶ Dark matter relic density:
 $\Omega_{\text{CDM}} h^2 = 0.1187 \pm 0.0017 \pm 0.0119$ (Planck '13)
- ▶ Direct detection limit from 225 live days of Xenon100 ('12)

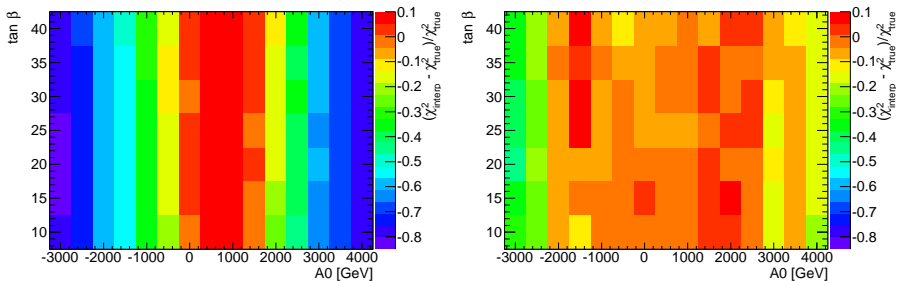
χ^2 contributions

At each parameter point \vec{P} calculate:

$$\chi^2 = \left(\vec{O}_{\text{meas}} - \vec{O}_{\text{pred}}(\vec{P}) \right)^T \text{cov}^{-1} \left(\vec{O}_{\text{meas}} - \vec{O}_{\text{pred}}(\vec{P}) \right) + \chi_{\text{limits}}^2$$



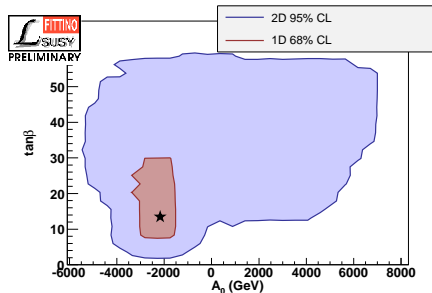
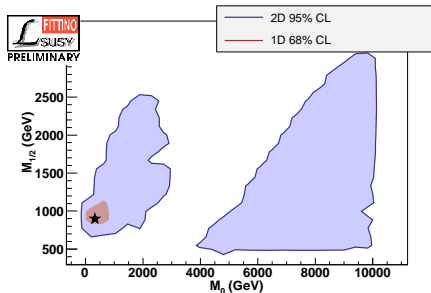
χ^2 contribution from ATLAS SUSY searches



- ▶ Depending on value of A_0 and $\tan \beta$ stop-production increases
- ▶ Along exclusion line we produced grids of correction factors in A_0 - $\tan \beta$

Preferred parameter space

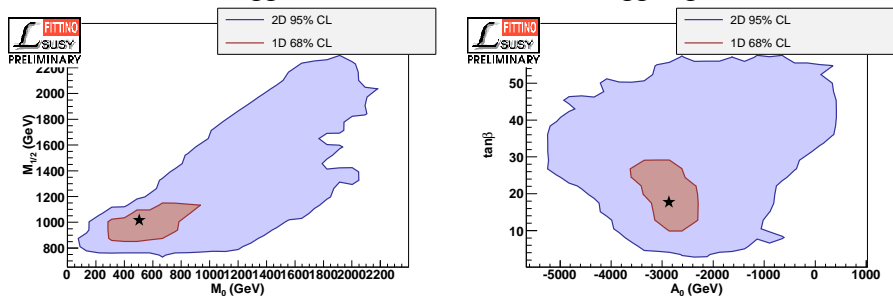
... with $m_h = (125.5 \pm 2 \pm 3)$ GeV but without Higgs rates



- ▶ $\chi^2/\text{ndf} = 13.6/9$
- ▶ FP / Higgs funnel region allowed at 2σ

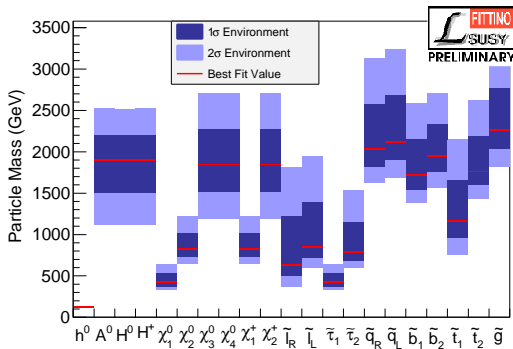
Preferred parameter space

... with mass and Higgs rates measurements via HiggsSignals



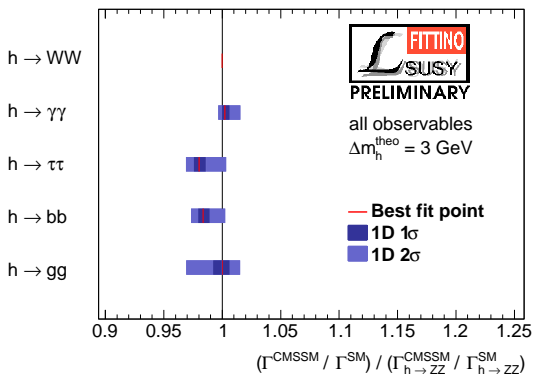
- ▶ $\chi^2/\text{ndf} = 49.6/59$: SM-like Higgs well described by CMSSM
- ▶ Fit quality seems to improve
- ▶ FP / Higgs funnel region disfavored at 2σ

Predicted mass spectrum



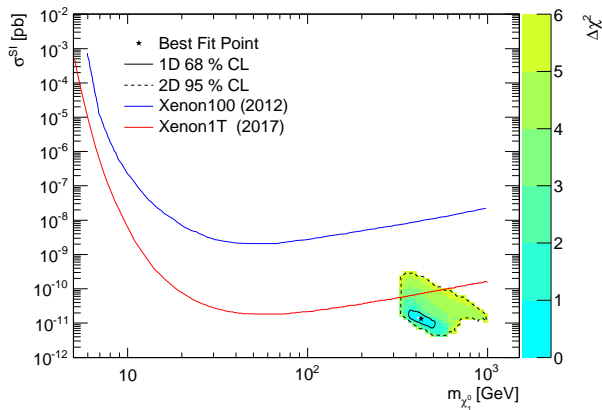
- ▶ Lower bound of 1 TeV on heavy Higgs bosons
- ▶ Relative light $\tilde{\tau}_1$ because of coannihilation region
- ▶ Relative light \tilde{t}_1 because of large stop mixing
- ▶ \tilde{q} and \tilde{g} masses at best fit point about 2 TeV

Predicted Higgs properties



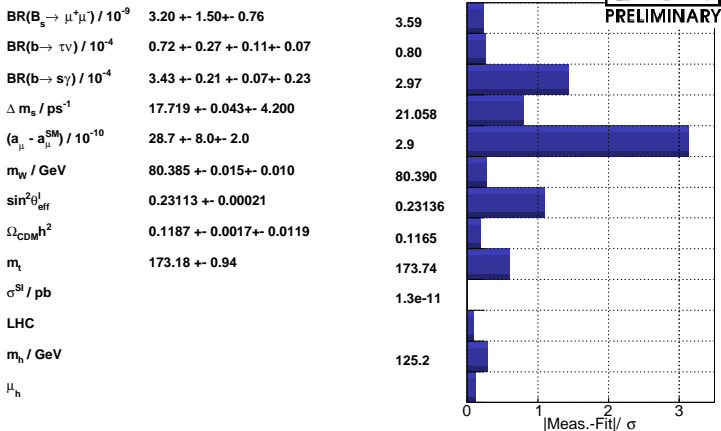
- ▶ SM-like Higgs boson
- ▶ not much room left for deviations

Predicted direct detection cross section



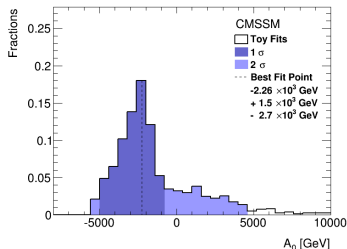
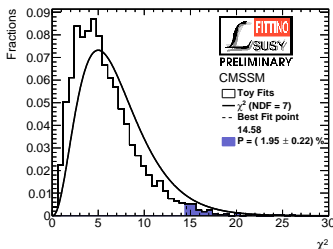
- ▶ Xenon 100 currently not sensitive
- ▶ Xenon1T will start to probe 2σ region

Agreement of observations with model predictions

 $M_0=504\text{GeV}, M_{1/2}=1016\text{GeV}, A_0=-2870\text{GeV}, m_t=174\text{GeV}, \tan\beta=18$


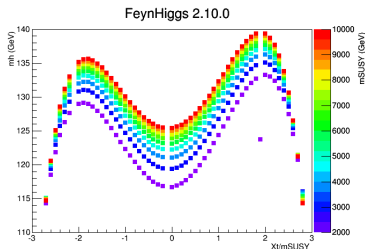
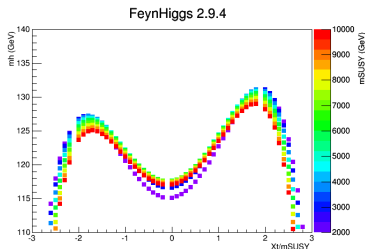
- ▶ SM like region of parameter space
- ▶ SUSY is losing its advantage in prediction of $g - 2$

Goodness of fit: Toy toys



- ▶ χ^2 function only guaranteed to be χ^2 -distributed if
 - ▶ uncertainties are Gaussian
 - ▶ observables depend linearly on parameters
- ▶ χ^2/ndf might overestimate the goodness of fit
 - ▶ here: proof of principle using (toy) toy fits
- ▶ bonus: toy distributions of parameters and observables

Why all results are preliminary



- ▶ Shown results use FeynHiggs 2.9.4
- ▶ The Higgs mass calculations used are not designed to work well for large values of m_{SUSY}
- ▶ This has been improved in FeynHiggs 2.10.0

Summary & Outlook

Summary:

- ▶ Best fit region of the CMSSM is very SM like
- ▶ CMSSM \sim SM + DM
- ▶ Difficult to distinguish them experimentally
- ▶ But: Relative light $\tilde{\tau}_1, \tilde{t}_1$ predicted
- ▶ To judge over the CMSSM, toy fits are needed

Outlook:

Stay tuned for results using

- ▶ Higgs mass calculation from FeynHiggs 2.10.0
- ▶ toy fits including Higgs rate measurements for calculation of \mathcal{P} -values